

Disks, black holes & gravitational waves

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CUNY (Grad Center, BMCC) & AMNH

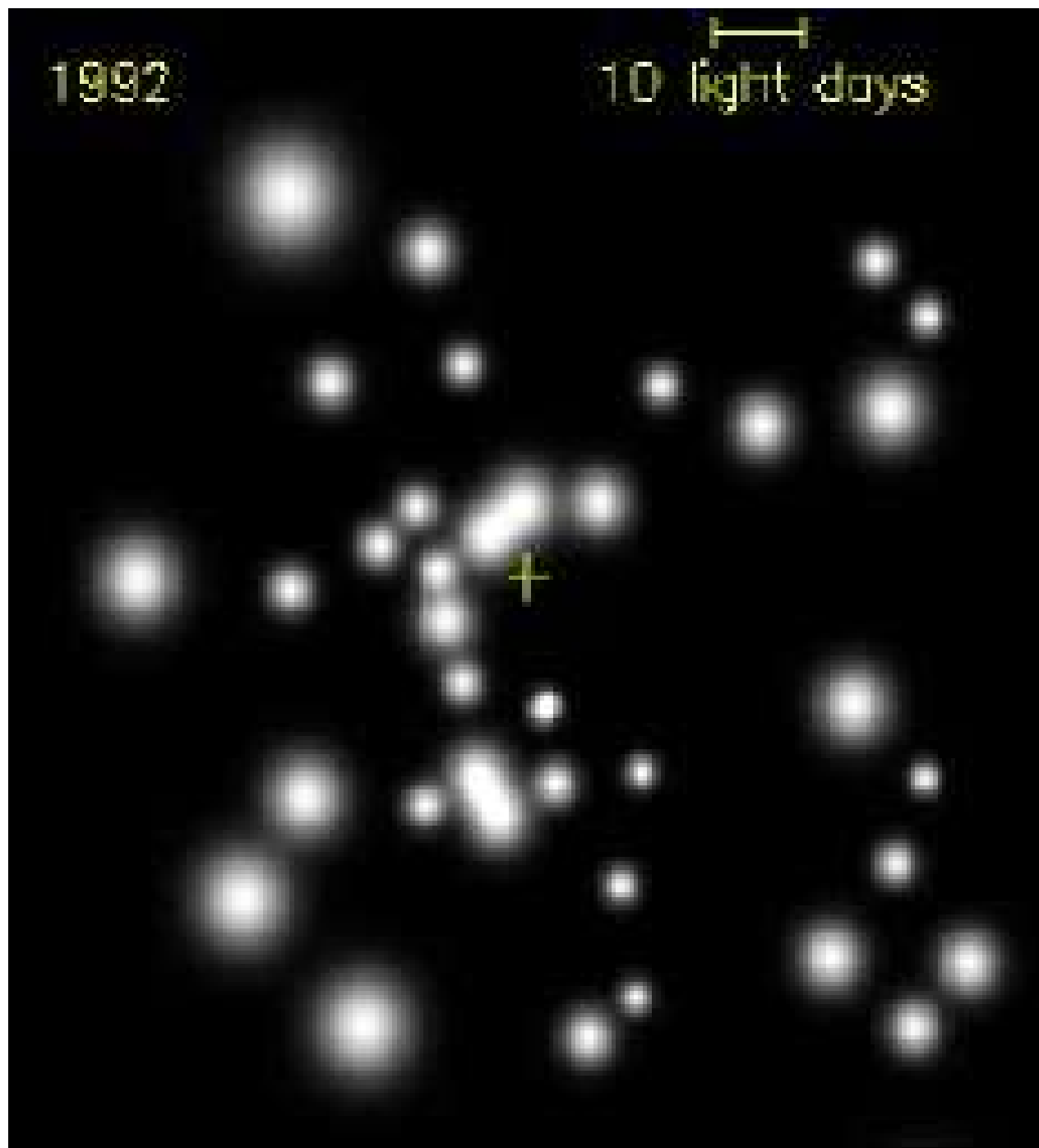
NYC: CUNY(Ford, Bellovary, O'Dowd, Minor), AMNH (Mac Low, Leigh, Shara), Columbia (Haiman)

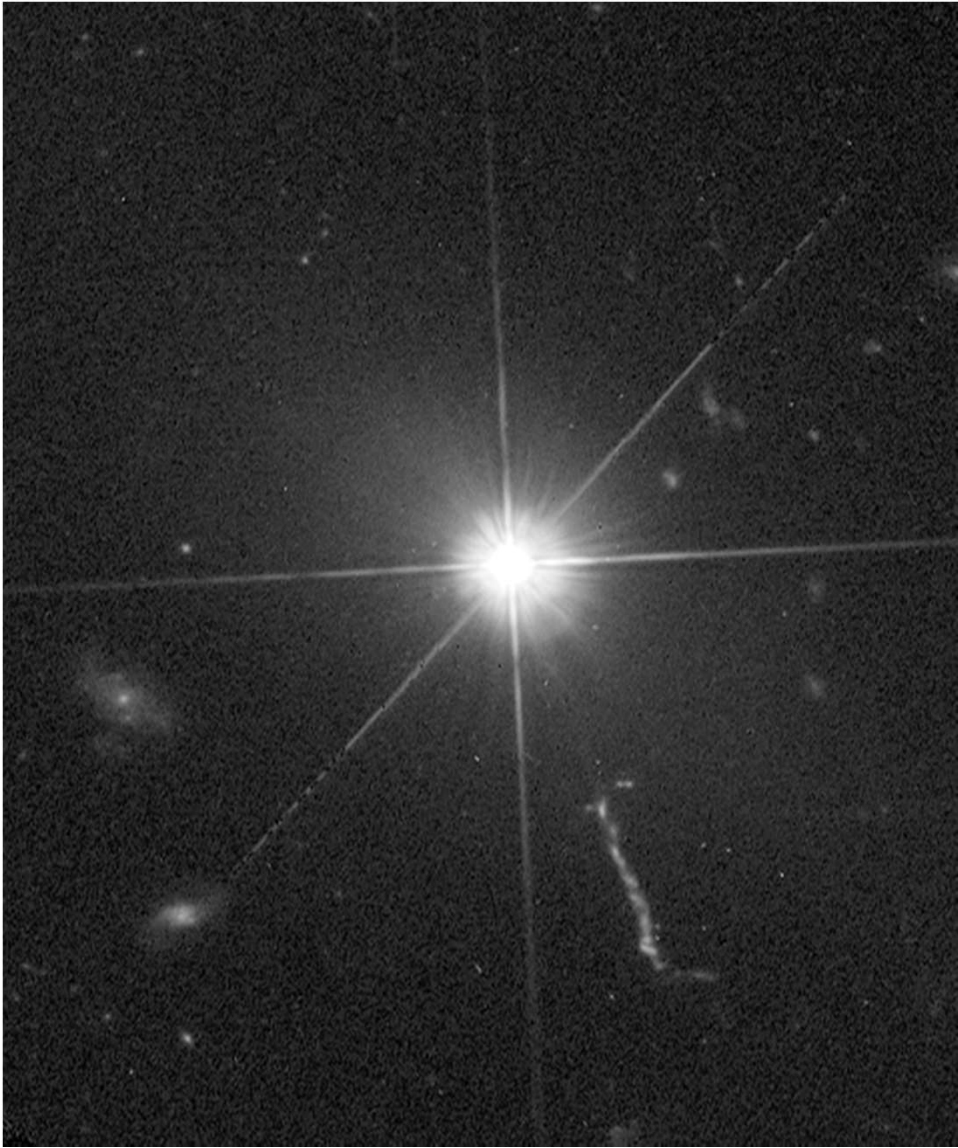
CalTech/JPL(Graham,Stern), STScI (Sivaramakrishnan)



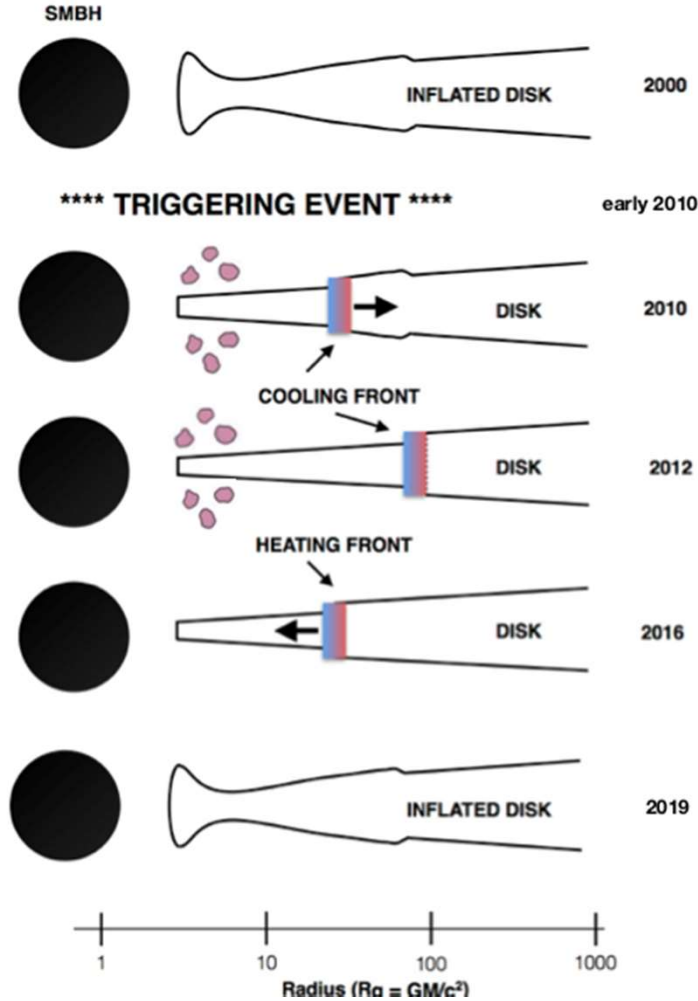
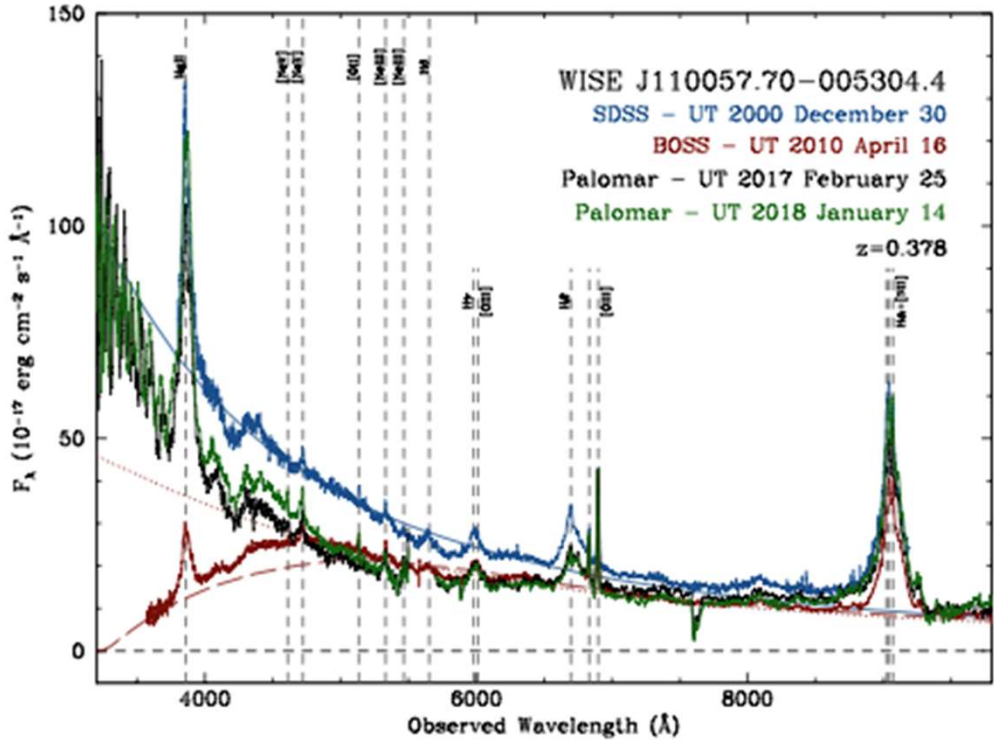
1992

10 light days





Rapidly changing disks



Stern+18; Ross+18; Graham+19

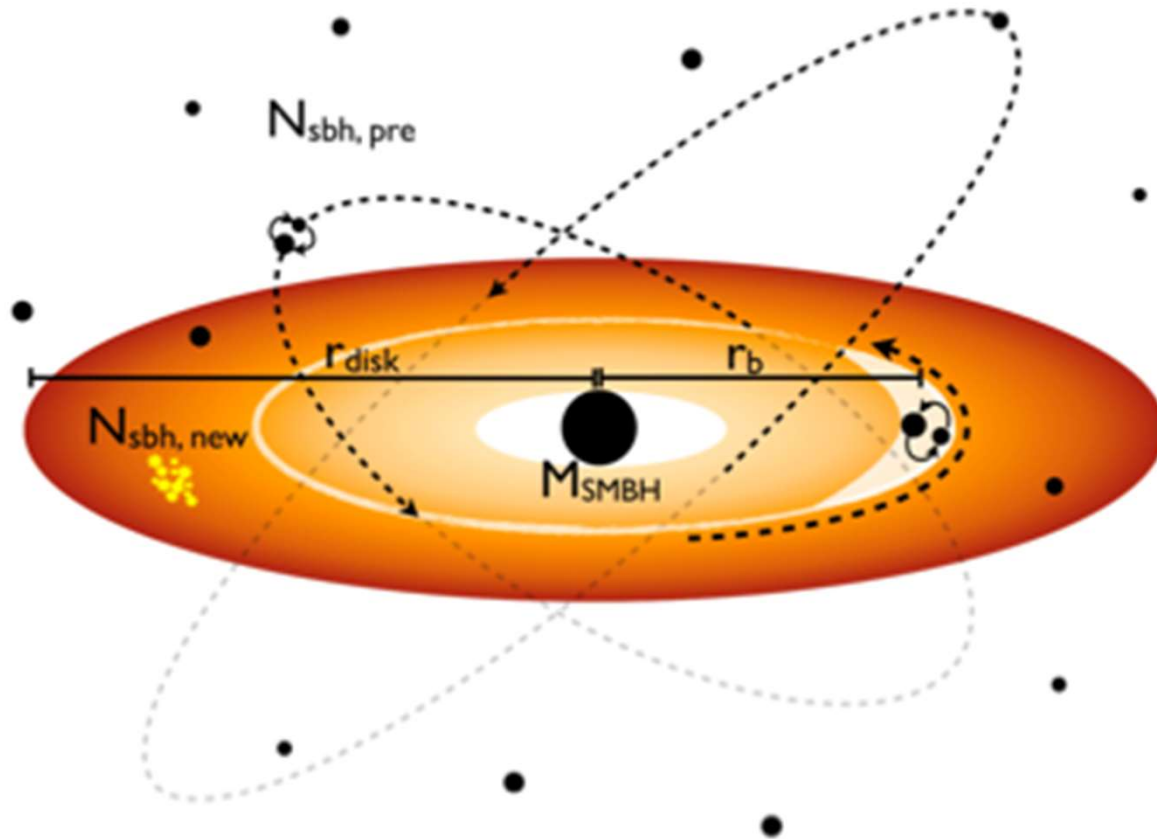
Galactic nuclei:

- Lots of stars & dead stars
- Densest population of BH in Universe (Hailey+18).

Add AGN gas disk:

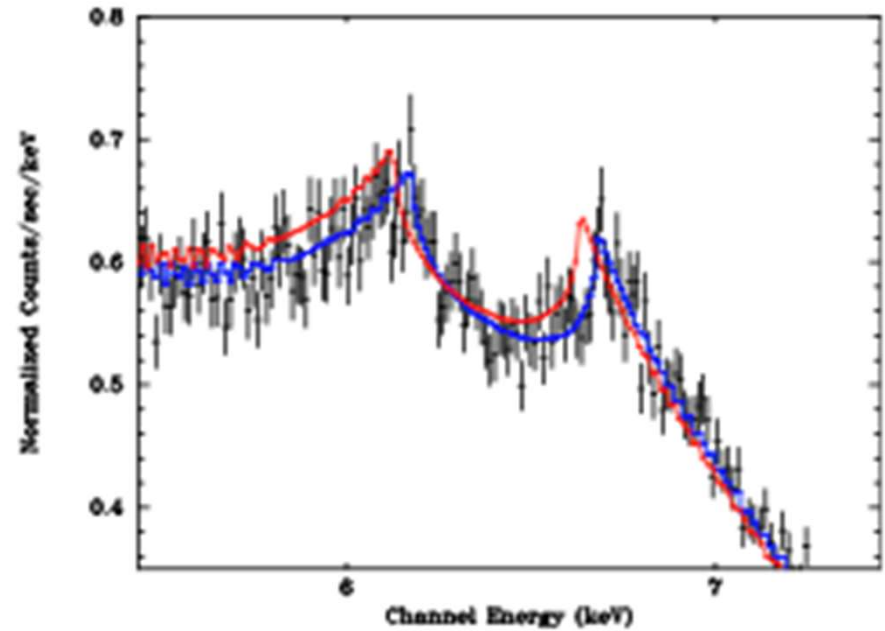
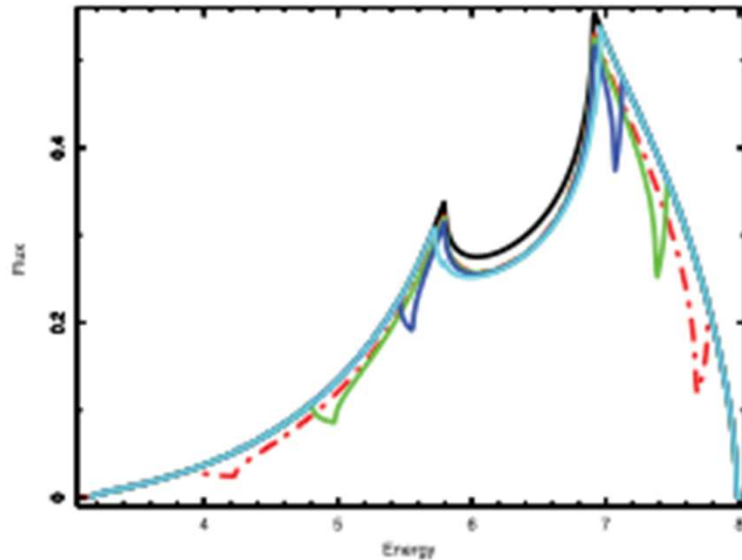
- Stars/BH end up in gas disk
- Migrate/collide & merge

A cartoon AGN



McKernan, Ford+2012, 2014=LIGO + LISA prediction
Bellovary+2016; McKernan+2018

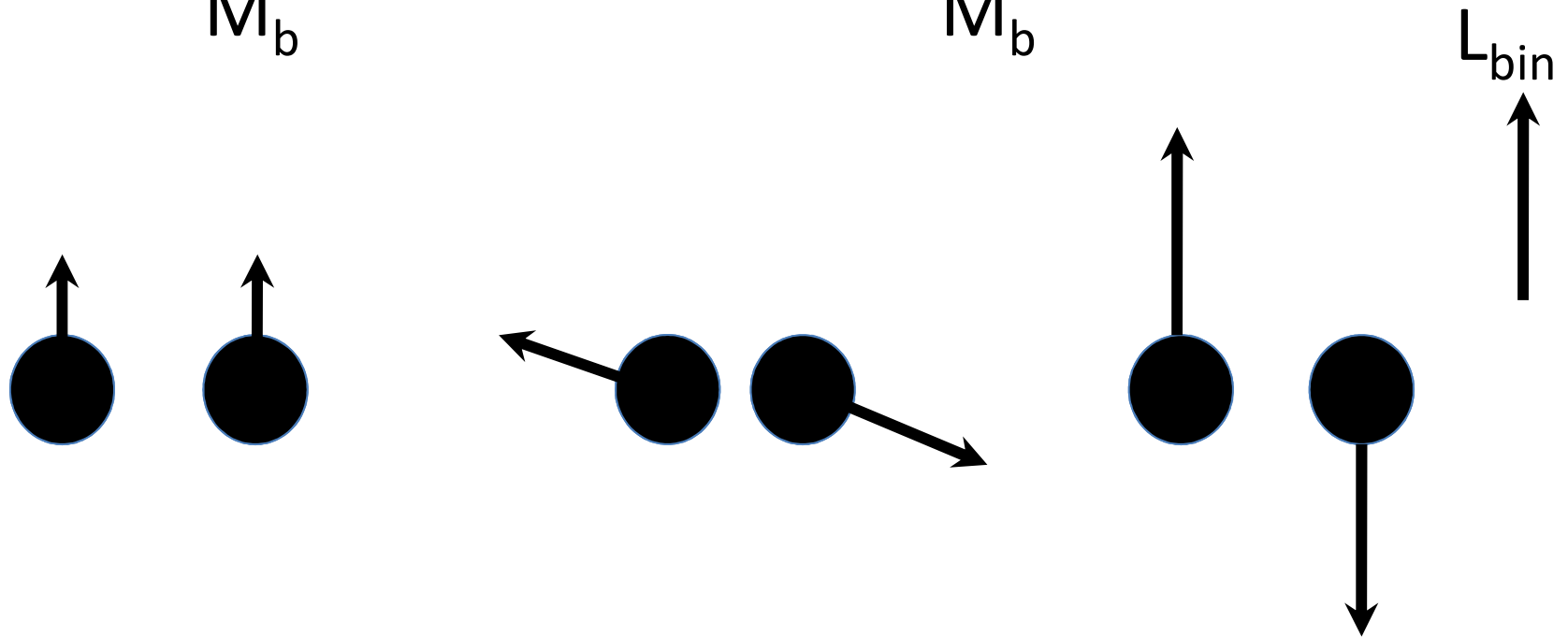
Relativistic lines as LIGO/LISA probes



McKernan, Ford, Kocsis & Haiman (2013); McKernan & Ford (2015)

Low $\chi_{\text{eff}} \neq$ low spin

$$\chi_{\text{eff}} = \frac{M_1}{M_b} |a_1| \cos\Theta_1 + \frac{M_2}{M_b} |a_2| \cos\Theta_2$$



χ_{eff} distribution prediction for LIGO O3 coming v. soon (McKernan & Ford 2018 in prep)